Rupture and Fault Zone Observatory

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The problem of earthquake generation remains unsolved because of its inherent complexity and the lack of direct observations within rupture zones (almost all data are low-passed-filtered elastic far-field radiation)

Key Questions

Pre rupture

-What processes produce the conditions that allow large earthquakes to occur, and how are they manifested in data? Improved Forecasting

During rupture

-What rheology governs brittle failure and permanent deformation within and around rupture zones? Improved physics of fault failure and GMPE

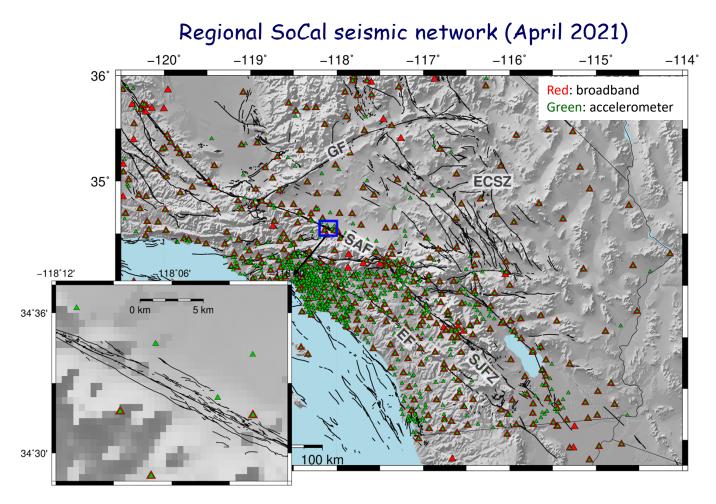
Post rupture

-What are the immediate and evolving post seismic processes (afterslip, viscoelastic, poroelastic) Improved quantification of strain energy budget on faults and crustal rheology



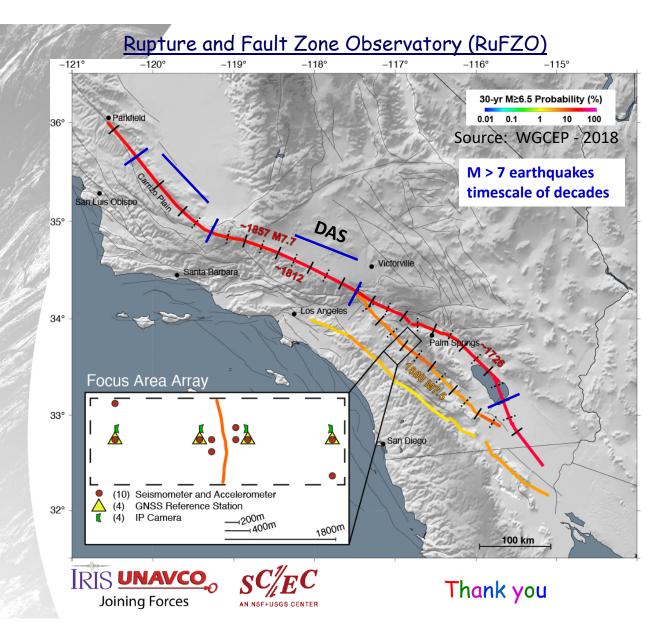






There are ~420 BB and ~1000 accelerometers, but very few stations within 1 km from main fault surfaces.

The lack of stations within-fault-zones is global. There are currently near-zero in-situ constraints on earthquake rupture processes!



<u>RuFZO data</u>

<u>Pre rupture</u>

-Evolving localization, temporal changes episodic local failures, other signals?

During rupture

-Full evolving dynamic strain field (shear and dilatational), dynamic/static strain/stress drops

-Dynamic rupture width and velocity, slip velocity (including space-time variations)

-Seismic energy flux (partitioning between radiation and dissipation)

-PGA/PGV/PGD within and near rupture zone (needed for improved GMPE)

-Robust Early Warning signals (including directivity)

Post rupture

Detailed transition from seismic to aseismic deformation, including volumetric components; evolving postseismic fields

<u>Between ruptures</u>

-High resolution 4D information around main faults; integration with regional data will improve event locations, focal mechanisms, slip inversions,

<u>Plus</u>

-Surprise discoveries and applications